

## Vicarious Calibration for MERIS 4<sup>th</sup> Reprocessing

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#### FRM4SOC

Options for future European satellite OCR vicarious adjustment infrastructure for the Sentinel-3 OLCI and Sentinel-2 MSI series Feb 21-23 2017, ESRIN

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- History: the MERIS 3rd reprocessing (M3RP) vicarious adjustment
- MERIS 4th reprocessing (M4RP) algorithmic evolutions
- Vicarious adjustment for M4RP
- Alternative methodology
- Results and validation



## MERIS 3rd reprocessing vicarious adjustment

- ATBD 2.24 M3RP (ACRI-ST 2011): "it is chosen to use the terminology of a vicarious adjustment, rather than calibration [...] The vicarious adjustment is not just a sensor calibration, but an adjustment of the whole system sensor+processing chain (in particular atmospheric correction). It should be updated at every change in the Level1 or Level2 ground segment"
- Adjustment on  $\rho_{GC}(\lambda) = \rho_{path}(\lambda) + t_d(\lambda)$ .  $\rho_w(\lambda)$  (TOA reflectance corrected for gaseous absorption, smile, and glint), not  $\rho_{TOA}(\lambda)$
- Uses the historical approach **decoupling VIS/NIR** for the assessment of vicarious gains (Franz et al. 2007, Bailey et al 2008), NASA ATBD modified in the NIR
- NIR adjustment over SIO/SPG clear waters
  - single-scattering approach (use aerosol reflectance and model retrieved at L2), different to NASA vicarious adjustment
  - assume two bands perfectly calibrated (709, 779 nm) to calibrate the other one with derived Angström exponent (865 nm)
  - the set of bands to fix/adjust is derived by a sensitivity analysis
- VIS vicarious adjustment using colocated in-situ measurements from BOUSSOLE and MOBY (open oceans, homogeneity of targets)
  - propagate in situ  $\rho_w^{IS}$  to TOA using satellite-retrieved atmosphere  $\rightarrow \rho_{GC}^{IS}(\lambda)$
  - gain =  $\rho_{GC}(\lambda) / \rho_{GC}^{IS}(\lambda)$



#### **MERIS 3rd final gains**

M3RP final gains computed from:

- 1) NIR gains: SIO/SPG
- 2) VIS gains: BOUSSOLE/MOBY

Uncertainties: standard deviations of per-matchup individual gains



2) VIS adjustment using in-situ colocated data

1) NIR adjustment using observations over SIO/SPG 709/775 fixed



- Changes in all steps of the processing: L1 calibration, preprocessing (gaseous transmissions, classification...), cloud, land and water branches including the possibility to process a pixel into different branches if ambiguity in the classification
- Water branch especially:
  - pressure adjustment scheme modified with an initial view on processing over high altitude lakes → see next
  - bright-pixel atmospheric correction (BPAC) evolutions, supposed to handle NIR adjustment → no NIR gains as BPAC performs spectral alignment similar to what is done in M3RP NIR gain computation
  - aerosol models now from Goddard Space Flight Center (GSFC) instead of the so-called Standard Aerosol Models (SAM)
- $\rightarrow$  no reason to obtain same set of gains as M3RP



### MERIS 4RP pressure adjustment change

- Pressure adjustment (both in M3RP and M4RP): adjust the signal, before atmospheric correction (AC), to a reference pressure over which radiative transfer model has provided look-up-tables (LUTs)
- AC: determine  $\rho_{path}(\lambda)$ ,  $t_d(\lambda)$  from aerosol models retrieved at 779/865 nm  $\rho_{GC}(\lambda) = \rho_{path}(\lambda) + t_d(\lambda) \cdot \rho_w(\lambda)$
- M3RP:
  - $\rho_{GC}(\lambda)$  at **local pressure** Ppix: corrected for **gas. abs., smile, glint**
  - $\rho_{path}(\lambda NIR)$  adjusted at Pref to perform AC (i.e. determine  $\rho_{path}(\lambda)$  over all spectrum) at reference pressure (1013 hPa)
  - $\rho_{path}(\lambda)$  retrieved at Pref then **deadjusted at Ppix**
- M4RP:
  - $\rho_{GC}(\lambda)$  at **local pressure** Ppix: corrected for **gas. abs., glint**
  - $\rho_{GC}^*(\lambda)$  at **reference pressure** P1: corrected for smile, Bodhaine (lat. dependency of Rayleigh), and pressure via equivalent Rayleigh optical thickness  $\tau_{RAY\_meas}(\lambda)$
  - AC performed on  $\rho_{GC}^*(\lambda)$  at **reference pressure** P1 closest to Ppix

→ the effect on VIS vicarious adjustment methodology is the necessary adjustment on  $\rho_{GC}^*(\lambda)$  in M4RP ( $\rho_{GC}(\lambda)$  at reference pressure)



#### MERIS 4RP: handle AC over high altitude lakes

- Goal: since lakes can be at high altitudes RTM LUTs have been built to model  $\rho_{GC}^*(\lambda)$  for reference pressures {1040, 1013, 970, 900, 800, 700} hPa
- Depending on water-body altitude one reference pressure level must used preferably
- A sensitivity analysis proved necessary to retrieve  $\rho_{GC}^*(\lambda)$  at two bracketing reference pressures P1 and P2:
  - $\rho_{GC_P1}^*(\lambda)$  and  $\rho_{GC_P2}^*(\lambda)$  with P1/P2 being the bracketing pressure levels
  - $\rho^*_{GC\_P1}(\lambda) \rightarrow \rho_{w\_P1}(\lambda)$
  - $\rho^*_{GC_P2}(\lambda) \rightarrow \rho_{W_P2}(\lambda)$
  - final  $\rho_w(\lambda)$  via interpolation using  $\tau_{RAY\_meas}(\lambda)$ ,  $\tau_{RAY\_P1}(\lambda)$ ,  $\tau_{RAY\_P2}(\lambda)$
  - $\rho_w(\lambda) = \alpha . \tau_{RAY\_meas}(\lambda) + \beta$
  - $\rho_{w_P1}(\lambda) = \alpha \cdot \tau_{RAY_P1}(\lambda) + \beta$
  - $\rho_{w_P2}(\lambda) = \alpha . \tau_{RAY_P2}(\lambda) + \beta$
- Vicarious adjustment in the VIS: perform over  $\rho^*_{GC_P1}(\lambda)$  and  $\rho^*_{GC_P2}(\lambda)$



- $\rightarrow$  Two propositions from QWG:
  - apply two set of gains separately for  $\rho^*_{GC_P1}(\lambda)$  and  $\rho^*_{GC_P2}(\lambda)$  then interpolate
  - compute one unique gain to apply on both  $\rho^*_{GC_P1}(\lambda)$  and  $\rho^*_{GC_P2}(\lambda)$
- Both solutions lead to same estimations of the gains over oceanic waters (BOUSSOLE / MOBY)
- Solutions may diverge **over high altitude targets**: gains determined over oceanic targets may not be transferable, but no possibility to assess (no in situ site available over high altitude lakes)

→ second solution kept by coherence with M3RP implementation of the gains through the L2 processor: one set of gains applied

$$\rho_{GC_P1}^{vic}(\lambda) = g(\lambda). \rho_{GC_P1}^*(\lambda) \text{ and } \rho_{GC_P2}^{vic}(\lambda) = g(\lambda). \rho_{GC_P2}^*(\lambda)$$



- latest BOUSSOLE and MOBY reprocessings since M3RP + more data
- latest MERIS processor
- individual gains computed per pixels within 5x5 matchups carefully filtered (no glint, no cloud, low AOT...)
- median gain computed per matchup  $\rho_{GC}(\lambda) / \rho_{GC}^{IS}(\lambda)$
- mean gain = weighted-average over all matchups
- uncertainties being:
  - $\sigma = \sqrt{\sigma^{sat^2} + \sigma^{IS^2}},$
  - with  $\sigma_{sat}$  std over the macropixel

• 
$$\sigma_{IS} = 5\%$$
 of  $\rho_w^{ISMI}$ 





#### MERIS 4th vicarious adjustment in the VIS

• Time series M3RP vs M4RP (example 490 nm):





#### M3RP vs M4RP vicarious gains



No reason to obtain same gains in M4RP as in M3RP:

- no gains in the NIR
- many processors changes, not only in the VIS



#### Alternative methodology

Perform a **zero adjustment** of the measuring system, i.e. offset, and a **span adjustment**, i.e. gain  $\rightarrow$  Vicarious Adjustment (VA)

- → apply gain and offset to  $\rho_{GC}^*(\lambda)$  to align  $\rho_w$  on  $\rho_{w, IS}$
- → proper zero or <u>no bias</u>: resolve a potential residual bias of the processor, corrected at AC input
- → proper representation of range for all potential measurements, i.e. quantity values being attributed to the measure, or <u>coverage</u> (≠ trueness though correcting systematic errors)
- $\rightarrow$  gain corrects instrumental errors as processor errors

 $\rho'_{gc}(\lambda) = \alpha(\lambda) \cdot \rho_{gc}(\lambda) + \beta(\lambda)$  where  $\alpha$  is the gain and  $\beta$  is the offset

#### **Recommended methodology:**

- atmospheric correction behaves like an « instrument » transfer function
- search for 'optimal' gains + offsets by
  - ${}^{\ensuremath{\mathscr{D}}}$  maximizing likelihood on estimated  $\hat{
    ho}_w$
  - taking matchups' uncertainties into account

(identification of ranges of "authorized" gains and offsets, with regards to samples' representativeness and to noise)

= better constrained adjustment but on the same line as M3RP 's



#### Alternative methodology: description

For each  $\alpha(\lambda)$  and  $\beta(\lambda)$  :

- apply gain+offset to  $\rho_{gc}(\lambda)$ :  $\rho'_{gc}(\lambda) = \alpha(\lambda) \cdot \rho^*_{GC}(\lambda) + \beta(\lambda)$
- compute  $\rho'_w(\lambda)$  such that  $\rho'_{gc}(\lambda) = \rho_{path}(\lambda) + t_d(\lambda) \cdot \rho'_w(\lambda)$
- compute mean of residuals:

$$R = \frac{1}{N} \sum \left( \frac{\rho_w^{ISME}(\lambda) - \rho'_w(\lambda)}{\sigma} \right)^2$$

•  $\sigma = \sqrt{\sigma^{sat^2} + \sigma^{IS^2}}$ , with  $\sigma_{sat}$  std over the macropixel,  $\sigma_{IS} = 5\%$  of  $\rho_w^{ISME}$ 

- $\rightarrow$  Find  $\hat{\alpha}$ ,  $\hat{\beta}$  minimizing *R*, band per band
  - as if there was no correlation between  $\rho_w$  at different wavelength, though there are because of the atmospheric transfer -



#### Alternative methodology (uncertainty)

Cramer-Rao bound for variance of an estimator (lower bound on the variance of unbiased estimators of a deterministic parameter: inverse of the Fisher information *I*)

$$\mathrm{Var}\left(\hat{ heta}
ight) \geq rac{1}{\mathcal{I}( heta)} = rac{1}{-\mathbb{E}\left[rac{\partial^2}{\partial heta^2}\ln f(X; heta)
ight]}$$

here we have an estimator of gain and offset  $\hat{\theta} = (\hat{\alpha}, \hat{\beta})$ with pdf as product of independent pdfs for each matchup

$$f = \prod \frac{1}{\sigma \sqrt{2\pi}} e^{-\frac{(\rho_w^{ISME}(\lambda) - \rho_w'(\lambda))^2}{2\sigma^2}}$$

where  $\rho'_w(\lambda)$  depends on  $\alpha(\lambda)$  and  $\beta(\lambda)$ 

- $\rightarrow$  constraints on Var( $\hat{\alpha}$ ) and Var( $\hat{\beta}$ ) (resp. on  $\sigma_{\hat{\alpha}}$  and  $\sigma_{\hat{\beta}}$ )
- → defines a domain where  $\alpha \pm \sigma_{\hat{\alpha}}$  and  $\beta \pm \sigma_{\hat{\beta}}$  can be considered as statistically acceptable solutions
- → such domain corresponds to acceptable values of the mean residual, upper bound = max R( $\alpha \pm \sigma_{\hat{\alpha}}$  and  $\beta \pm \sigma_{\hat{\beta}}$ )

#### Alternative methodology (result)



Domain of R with  $Var(\hat{\theta}) < boundary$  as function of  $\alpha(\lambda)$  and  $\beta(\lambda)$ :



 $\rightarrow$  gain with offset=0 is among solutions (same at all wavelengths)

 $\rightarrow$  can be a solution based on a priori knowledge that offset=0



### Alternative methodology: comparisons with M4RP

Domain of R with  $Var(\hat{\theta}) < boundary$  as function of  $\alpha(\lambda)$  and  $\beta(\lambda)$ , valid domain for both BOUSSOLE and MOBY





Validation

# Comparison nogains / M4RP gains: removing bias differences to IS, 412 nm





#### Comparison M3RP / M4RP (RPD): 412



AAOT

Algarve

AbuAlBukhoosh

BOUSSOLE

BristollrishSea

CaliforniaCurrent

Validation



#### Validation



Less dispersion in M4RP

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#### Validation



Less dispersion in M4RP

MERMAID data



- ATBD 2.24 M3RP: "vicarious adjustment [...] should be updated at every change in the Level1 or Level2 ground segment"
- $\rightarrow$  that is what is done for M4RP
- M4RP evolutions

→ adaptation and update of the vicarious adjustment, leading to very different vicarious gains profiles between M3RP and M4RP

→ M4RP: less bias and dispersion in comparison with MERMAID data

 many algorithmic changes in M4RP reprocessing chain are already transferred to S3 OLCI operational processing

 $\rightarrow$  some aspects of the vicarious adjustment can be transferred to OLCI



#### **Expectations from FRM4SOC**

- provide total uncertainty budget to better constrain the gains computation
- systematically provide ancillary parameters (e.g. wind, cloudiness, wave height...) + AOT / Angström in the NIR to better constrain the retrieval and the analysis
- in situ data deliveries: common naming conventions and quantities would be nice



#### THANK YOU !

